

Stellar Explosions: Core-collapse Supernovae and Nuclear Theory

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FRIB Theory Alliance Meeting 1 April 2016



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We've solved the problem once and for all!!



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Sorry.... April Fool's!

Nucleosynthesis from CCSNe

TA	Big Bang nucleosynthesis spallation evolved giant stars α-elements iron group elements									α-rich freeze-out, vp-process, weak s-process?									
1 H										s-process light neutron-capture primary process									
3 Li										r-process			6 C	7 N	8 O	9 F	10 Ne		
11 Na 22.990	12 Mg 24.312											13 Al 26,982	14 Si 28.086	15 P 30,974	16 S 32.064	17 Cl 35,453	18 Ar 39.948		
19 K 39,102	20 Ca 40.08	21 Sc 44,956	22 Ti 47.88	23 V 50,942	24 Cr 51,996	25 Mn 54,938	26 Fe 55,847	27 Co 58,933	28 Ni 58,69	29 Cu 63,54	30 Zn 65,37	31 Ga 69.72	32 Ge 72,59	33 As 74,922	34 Se 78,96	35 Br 79,909	36 Kr \$3,80		
87 Rb 85.47	38 Sr 87,62	39 Y 88,905	40 Zr 91,22	41 Nb 92,906	42 Mo 95	43 Tc (99)	44 Ru 101,07	45 Rh 102.91	46 Pd 106.42	47 Ag 107,87	46 Cd 112.40	49 In 114,82	50 Sn 118.69	51 Sb 121,75	52 Te 127,60	53 I 126,90	54 Xe 131,30		
55 CS 132,91	56 Ba 137,34	57 La 138,91	72 Hf 178.49	73 Ta 180,95	74 W 183.85	75 Re 186,2	76 OS 190,2	77 Ir 192,2	78 Pt 195.09	79 Au 196,97	80 Hg 204.59	81 TI 204,38	82 Pb 209,17	83 Bi 208,98	84 Po (210)	85 At (210)	86 Rn (222)		
87 Fr (223)	88 Ra (226)	89 Ac (227)																	
		1	58 Ce	⁵⁹ Pr	⁶⁰ Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy 162 50	67 Ho 164.93	68 Er	69 Tm	70 Yb	71 Lu			
			90 Th 232.04	91 Pa (231)	92 U 238.03	93 Np (237)	94 Pu (242)	95 Am (243)	96 Cm (247)	97 Bk (249)	98 Cf (251)	99 Es (254)	100 Fm (253)	101 Md (256)	102 No (253)	103 Lr (257)			

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CCSNe dominate the chemical evolution of the universe

A. Frebel, MIT





Nuclear Physics Labs

- Matter at most extreme densities, temperatures, isospin
- Produce most elements in Nature
- Complimentary to experiment at, e.g., NSCL, FRIB, JLAB, ATLAS, RHIC, GSI, TRIUMF,...
- Neutrino and gravitational wave signals encode info about nuclear EOS
- Test BSM physics of neutrinos





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2000 km

Fe

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10⁵⁷ neutrinos released!!



 $Y_e \sim 0.27$











50 km

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$\mathrm{Fe} \rightarrow 4 \mathrm{He}$

M

Shock stalls... What revives it?? => The CCSN "Problem"



Dense Matter Equation of State

- Can impact:
 - Explosion (Marek et al. 2009; SMC 2013; Suwa et al. 2013)
 - Gravitational wave emission (Marek et al. 2009; Mueller et al. 2013)
 - Neutrino emission (O'Connor & Ott 2013)



Fischer et al. 2014

S.M. Couch



3D CCSN Simulations





Time=0.299 s



200 km



3D CCSN Simulations





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200 km



3D with Full Nu Transport Janka, Melson, & Summa (2016)

3D explosions required:

- low-mass progenitor
- (unphysically) large strangeness correction
- rapid rotation





Results from Oak Ridge Group



II Nu Transport (2015), Lentz et al. (2015)



3D explodes later than 2D Only 2° resolution...



Resolution Dependence

D. Radice, C. Ott, SMC, et al., ApJ, 820, 76

- Analytic accretion shock IC's
- Inner boundary
- Lightbulb heating/cooling
- Fixed-metric GR



3D Slices





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3D Slices





3(D) steps forward, 1(D) step back?

- physics wrong...
- Possibilities:
 - Progenitor structure
 - MHD/rotation
 - Behavior of turbulence/low-resolution O

 - Equation of state
 - Nuclear physics (i.e., strangeness)

Success of explosions in 2D may not be recovered in 3D... We must be missing physics, or getting the

• Neutrino effects (i.e., oscillation, x-sections, sterile)



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CCSN is an Initial Value Problem ...And we have problems with our initial values

- All stars rotate
- All stars have magnetic fields
- binaries

• Stars are not (perfectly) spherical

>50% of SN progenitors in interacting



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Unperturbed



Progenitor Asphericity in 3D SMC & Ott (2013, 2015); SMC et al. (2015)

Time=0.251 s



200 km

Perturbed





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J. Lippuner, L. Roberts

JINA Online Seminar, 29 January 2016

Connecting to Observation

With an explosion model:

- **Gravitational Waves**
- Neutrinos
- NS/BH mass distributions
- Nucleosynthesis
- Light curves/spectra
- Nuclear data











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JINA Online Seminar, 29 January 2016

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Toward predictive CCSN theory

- Modern high-performance computing revolutionizing CCSN theory
- CCSN mechanism depends critically on nuclear theory input
- Understanding sensitivity to nuclear physics requires robust explosion model
- Goal of predictive CCSN theory within reach!

